

FN-372 0402.000 (Submitted to Particle Accelerators)

PHASE DYNAMICS NEAR TRANSITION ENERGY IN THE FNAL MAIN RING

K. Takayama

July 1983

Phase Dynamics near Transition Energy in the FNAL Main Ring

### Ken TAKAYAMA

Fermi National Accelerator Laboratory , Batavia, ILL 60510, USA

# Contents

- 1. Introduction
- 2. Difference and Differential Equations for Acceleration
- 3. Hamiltonian Formalism
- 4. Linear Motion
- 5. Nonlinear Motion
- 6. Emittance Blowup at Transition
- 7. Application
- 8. Conclusion

Acknowledgments

Appendixes

References

#### Abstract

The phase dynamics of smallamplitude synchrotron oscillations in the vicinity of the transition energy is discussed with kinematic nonlinearities included. We introduce synchrotron amplitude function analogous to the betatron amplitude function and solve analytically the time evolution of bunch shapes, where the kinematic nonlinearities result in unsymmetric bunch shapes. In addition, the above synchrotron oscillation is singular at transition crossing because of the kinematic nonlinearity. From this simple fact, we identify an inherent source of bunch diffusion. A method for estimating its size is presented. When this theory is applied to the case of the FNAL Main Ring, the predictions are in good agreement with numerical simulations and are not inconsistent with experimental results.

#### § 1. Introduction

The effect of nonlinear kinematic terms /1,2,3,4/ is studied for energies below and above transition energy. These nonlinear kinematic terms are stronger the narrower the bunches. The momentum height of the bunch passes through a maximum at transition and the kinematic terms therefore have a maximum at transition. They can ditort the particle orbits in different ways, and they may lead to longitudinal emittance blowup.

In recent experiments in the FNAL Main Ring, bunch lengths were measured at two energies, 14 GeV and 19.7 GeV, below and above transition, at an average intensity of 2.6\*1010 protons per bunch (total Main Ring intensity 2.8\*1013 protons per cycle). Values for the longitudinal emittance at the two energies have derived from these measurements. The results were 0.22 eV-sec at 14 GeV and 0.28 eV-sec at 19.7 GeV, indicating an emittance increase in the region of transition. Bunch lengths were also measured at transition (17.6 GeV), where they become very narrow (about 2.5 nsec). Furthermore, in order to clarify the reasons which lead to this longitudinal emittance blowup at transition crossing, many extensive computer simulations been performed independently by several people, including the present author. The simulation results, which strongly imply that the effects of the nonlinear kinematic term are large, are surprisingly consistent with the measurements.

It is the purpose of this paper to calculate the effects of

nonlinear kinematic terms in the range around the transition energy and compare to results of computer simulations and real machine studies.

This paper is divided in four main parts: In the first part (§ 2,3) we derive difference equations for acceleration in an explicit form and transform them into an differential form. which enables us to construct a Hamiltonian formulation for longitudinal motion. Here we shall restrict ourselves to small amplitude oscillation. addition, only the lowest-order nonlinear Τn kinematic term will be retained in this formulation. part (§ 4), introducing the notion of a synchrotron amplitude function, we construct the "linear classical theory" of transition in a form analogous to betatron oscillation, where the nonlinear kinematic term is neglected. In the third part (§ 5), using the perturbation theory, we calculate increments of the longitudinal emittance as an effect of the nonlinear kinematic term on linear motion. We identify this effect as a reason of the unsymmetric bunch shape just at transition that has been recognized in the computer simulations. In the fourth part (§6), from a general point of view with respect time-reversability, it will be shown that only such nonlinear kinematic terms can accumulate to give net effects over transition crossing. Finally, a theoretical for formula for the emittance blowup ratio will be presented.

In the present discussions, effects of longitudinal space-charge forces/6/ and timing error of the phase-jump at

transition /7/ are not included, because the former is negligible at least for the present situation of the FNAL Main Ring and exact information about the latter has not been obtained.

## § 2. Difference and Differential Equations for Acceleration

The theory of longitudinal phase motion, describing the energy and phase oscillatios that occur when a particle passes repeatively through one or more "accelerating cavities" situated at localized points around the accelerator ring, is well known. Since the oscillations normally are at a relatively low frequency, it is often legitimate as well as convenient to analyze them theoreticall with differential equations derived by spreading the accelerating field uniformly around the orbit. In reality, the energy changes experienced by a particle are better represented by difference equations and depend on the sine of the electrical phase angle  $\varphi$  at which the particle traverses the cavity. The corresponding equations of motion are therefore both nonlinear and discrete.

We consider here the case of synchrotron oscillations during acceleration stage. To obtain the actual transformation, we consider a short cavity system operating at a harmonic number h, an angular frequency  $\omega_{r,j}(t)$  and a peak voltage V(t). We assume that  $\omega_{r,j}(t)$  and V(t) are independently controlled during acceleration. The quantities denoted by  $E^n$  and  $\Phi^n$  are,

respectively, the energy and the electrical phase angle with which a particle enters the cavity at the time of transit. Then the nonlinear transformation may be written in the form

$$E^{n+1} = E^n + eV(n) \sin \varphi^n. \qquad (2-1a)$$

$$\Phi^{n+1} = \left\{ \frac{\omega_{rf}(n+1)}{\omega_{rf}(n)} \Phi^{n} + \omega_{rf}(n+1) \cdot \frac{2\pi}{\omega(E^{n+1})} \right\}, \quad (2-1b)$$

where  $eV(n)\sin \Phi^n$  is the energy gain at the n-th transit and the revolution period is described in the form

$$\frac{2\pi}{\omega(E^{n+1})} = \frac{C_0 (1 + \alpha_p z^{n+1})}{c \left[1 - (m_0 c^2 / E^{n+1})^2\right]^{\frac{1}{2}}} \left(z^{\frac{n+1}{2}} \frac{E^{n+1} - E_s^{n+1}}{\beta^2(E_s^{n+1}) E_s^{n+1}}\right), (2-2)$$

where  $C_0$  is the length of the closed orbit corresponding to the synchronous energy  $E_S$ , c is the velocity of light,  $m_0c^2$  is the proton rest energy, z is the momentum deviation from the synchronous momentum, and  $\alpha_P$  is the momentum compaction factor. Since we assume the guiding magnetic field of H type, the betatron acceleration can be neglected here. The synchronous particle is defined by the equations

$$E_s^{n+1} = E_s^n + \Delta(n),$$
 (2-3a)

$$\phi_s^{n+1} = \left\{ \frac{\omega_{r_4(n+1)}}{\omega_{r_4(n)}} \phi_s^n + \omega_{r_4(n+1)} \cdot \frac{2\pi}{\omega(E_s^{n+1})} \right\}, \qquad (2-3b)$$

where

$$\Delta(n) = eV(n) \sin \phi_s^n, \qquad (2-3e)$$

$$\omega_{rf}(n) \simeq h \omega(E_s^n) + \frac{\omega(E_s^n)}{2\pi} \left[ \sin^{-1} \frac{\Delta(n)}{eV(n)} - \sin^{-1} \frac{\Delta(n-1)}{eV(n-1)} \right].$$
 (2-3d)

Note that  $\Delta(n)$  is determined by the change in the external guide field B(t). Now the momentum compaction factor  $d_{p}/4/$  may be written in the form

$$\alpha_{p} = \alpha^{(0)} + \alpha^{(1)} (2^{n+1})^{1} + \alpha^{(2)} (2^{n+1})^{2} + O((2^{n+1})^{3}), \quad (2-4)$$

where  $O((\mathbf{Z}^{n+1})^3)$  is the Landau symbol. Expanding the right-hand side of Eq.(2-2) with respect to  $\mathbf{Z}^{n+1}$ , we have the expression

$$\frac{2\pi}{\omega(E^{nrt})} = \frac{C_0}{c\beta(E_s^{nrt})} \left[ 1 + \gamma^{(0)}_{(nrt)} Z^{nrt} + \gamma^{(1)}_{(2nrt)} (Z^{nrt})^2 + \gamma^{(2)}_{(2nrt)} (Z^{nrt})^3 + \dots \right]_{(2-5)}$$

$$\frac{C_0}{c\beta(E_s^{nrt})} = \frac{2\pi}{\omega(E_s^{nrt})} = \frac{2\pi h}{\omega_{rf}(nrt)}$$

$$\gamma^{(0)}(n+1) = \alpha^{(0)} - 1/\gamma^2(E_s^{n+1}) \quad (\alpha^{(0)} = 1/\gamma_T^2), \quad (2-6a)$$

$$7^{(1)} (n+1) = \alpha^{(1)} - \frac{7^{(0)}(n+1)}{r^{2}(E_{s}^{n+1})} + \frac{3\beta^{2}(E_{s}^{n+1})}{2r^{2}(E_{s}^{n+1})} = \alpha^{(1)} + \frac{3\beta^{2}(E_{s}^{n+1})}{2r^{2}(E_{s}^{n+1})} (2-6b)$$

$$7^{(2)} (n+1) = \alpha^{(2)} - \frac{2\beta^{2}(E_{s}^{n+1})}{r^{2}(E_{s}^{n+1})} + \frac{\alpha^{(1)}}{r^{2}(E_{s}^{n+1})} + \frac{7^{(0)}(n+1)}{r^{2}(E_{s}^{n+1})} + \frac{\beta^{2}(E_{s}^{n+1})}{2r^{2}(E_{s}^{n+1})} [3\alpha^{(0)} - \frac{5}{r^{2}(E_{s}^{n+1})}]$$

$$= \alpha^{(2)} - \frac{2\beta^{2}(E_{s}^{n+1})}{r^{2}(E_{s}^{n+1})}$$

We note that all particle simulations for a real acceleration mode stated in the Introduction have been performed by following the difference equations (2-1a) and (2-1b). The pair E and  $\varphi$  is recognized not to be canonical because of the time dependence of the frequency. We are interested in small-amplitude oscillations around the synchronous point ( $\varphi_s^n$ ,  $E_s^n$ ) as a guiding center. Setting

$$\epsilon^n = \epsilon^n - \epsilon_s^n \qquad (2-7a)$$

$$\chi^{n} = \varphi^{n} - \varphi^{n}_{s}, \qquad (2-7b)$$

we may write the difference equations for a small amplitude oscillation as

$$\epsilon^{n+1} = \epsilon^{n} + eV(n) \left( \sin \phi^{n} - \sin \phi_{s}^{n} \right) 
= \epsilon^{n} + eV(n) \left( \cos \phi_{s}^{n} \chi^{n} \right)$$
(2-8a)

$$\chi^{n+1} = \left\{ \frac{\omega_{rf}(n+1)}{\omega_{rf}(n)} \chi^{n} + \left[ 2\pi h + \sin^{-1} \frac{\Delta(n+1)}{eV(n+1)} - \sin^{-1} \frac{\Delta(n)}{eV(n)} \right] \right\}^{(2-8b)}$$

$$\left[ \frac{\gamma^{0}}{m}(n+1) \in \mathbb{R}^{n+1} + \frac{\gamma^{0}}{m}(n+1) \left( e^{n+1} \right) \right]^{(2-8b)}$$

$$\frac{7^{(0)}(n+1)}{\beta^2(E_s^{n+1})E_s^{n+1}},$$
(2-9a)

$$\overline{\gamma}^{(1)}(n+1) = \frac{\gamma^{(1)}(n+1)}{[\beta^{1}(E_{s}^{n+1})E_{s}^{n+1}]^{2}}$$
(2-9b)

To write down the difference equations in the form of exact differential ones we may use a  $\delta$  -function:

$$\dot{\epsilon} = \frac{2\pi \, e \, V(t) \, \cos \varphi_s(t)}{T_s(t)} \, \chi \, \delta_{2\pi}(t'), \qquad (2-10a)$$

$$\dot{\chi} = \frac{\dot{\omega}_{r,j}(t)}{\omega_{r,j}(t)} \chi + \frac{\left[2\pi h + T_s(t)\dot{\varphi}_s(t)\right]}{T_s(t)} \cdot \left[\overline{\gamma^o}_{(t)} \epsilon + \overline{\gamma^o}_{(t)} \epsilon\right]_{(2-10b)}$$

where  $T_{S}(t)$  is the period of the synchronous particle and one iteration of the mapping. Here  $t'=\Omega(t)t+t'$ ;  $\Omega(t)=2\pi/T_{S}(t)$  and the S -function of period  $2\pi$  is given by the Fourier expansion

$$S_{2\pi}(t') = \frac{1}{2\pi} \left( 1 + 2 \sum_{n=1}^{\infty} (os nt') \right).$$
 (2-11)

After neglecting rapidly oscillating terms in Eq. (2-10a), we have

$$\dot{\epsilon} = \frac{eV(t)(\cos \Phi_s(t))}{T_s(t)} \chi, \qquad (2-12a)$$

$$\dot{\chi} = \frac{\dot{\omega}_{rf}}{\omega_{rf}} \chi + \frac{\left[2\pi h + T_s(t) \dot{\varphi}_s\right] \left[7^{(0)} \dot{\epsilon} + 7^{(1)} \dot{\epsilon}^2\right]}{T_s(t)}.$$
 (2-12b)

## 3. Hamiltonian Formalism

Under the assumption that the damping term in Eq.(2-12b) is negligible in the short period of transition crossing, we can construct a Hamiltonian formalism for small-amplitude oscillations. Neglecting the damping term, we find  $(\mathcal{X}, \mathcal{E})$  to be a canonical pair that yields the Hamiltonian

$$H(\chi, \epsilon; t) = \frac{1}{T_{s}(t)} \left[ 2\pi h + T_{s}(t) \dot{\phi}_{s}(t) \right] \left[ \frac{1}{2} \overline{\eta^{(0)}}_{(t)} \dot{\epsilon}^{2} + \frac{1}{3} \overline{\eta^{(0)}}_{(t)} \dot{\epsilon}^{3} \right] - \frac{e V(t) (\cos \phi_{s}(t))}{2 T_{s}(t)} \chi^{2}.$$
(3-1)

We shall assume that the synchronous phase angle  $\phi_s$  jumps discontinuously at t=0 in such a way that  $\sin\phi_s$  is constant and

$$sgn(cos \phi_s) = - sgn(t)$$

We shall measure t from this instant. Now we introduce a scale change of the independent variable t by

$$\gamma(t) = \int_0^t \left[ -\frac{eV(t)(os\phi_s(t))}{T_s(t)} \right] dt$$
 (3-2)

Note that the new independent variable  $\gamma$  has dimension of energy. With the above origin for the time t,  $\gamma$  is always positive at all t, namely,

 $t \rightarrow 0$ -(approaching transition), then  $\Upsilon \rightarrow 0_+$ ,

 $0+ \rightarrow t$  (leaving transition), then  $0+ \rightarrow \gamma$ .

Such a scale change yields the new Hamiltonian

$$H'(x, \epsilon; x) = -\frac{\left[2\pi h + T_{s}(t)\phi_{s}(t)\right]}{eV(t)(os\phi_{s}(t))} \left[\frac{1}{2} \overline{\gamma''(t)} \epsilon^{2} + \frac{1}{3} \overline{\gamma''(t)} \epsilon^{3}\right] + \frac{1}{2} x^{2}.$$
(3-3)

For later convenience, let change the notation of the canonical variables to

$$\chi = p$$
(3-4a)

$$\epsilon = -x$$
. (3-4b)

Thus we obtain the Hamiltonian in the simple form

$$H''(x, p; \gamma) = \frac{1}{2}p^2 + \frac{1}{2}\lambda_0(\gamma)\chi^2 - \frac{1}{3}\lambda_1(\gamma)\chi^3, \qquad (3-5)$$

where

$$\lambda_o(t) = -\frac{\left[2\pi h + T_s(t)\dot{\phi}_s\right]}{eV(t)\cos\phi_s(t)}\frac{7^{o}(t)}{7^{o}(t)} = -\frac{\left[2\pi h + T_s(t)\dot{\phi}_s\right]\left[\alpha^{(e)} 1/r_s^2(t)\right]}{\beta_s^2(t)E_s(t)eV(t)\left(\cos\phi_s(t)\right)}, \quad (3-6a)$$

$$\lambda_{s}(\tau) = -\frac{\left[2\pi h + T_{s}(t)\dot{\phi}_{s}\right]}{eV(t)\cos\phi_{s}(t)}\overline{\gamma^{(i)}(t)} = -\frac{\left[2\pi h + T_{s}(t)\dot{\phi}_{s}\right]\left[a^{(i)} + \frac{3\beta_{s}^{2}(t)}{2F_{s}^{2}(t)}\right]}{\left[\beta_{s}^{2}(t)E_{s}(t)\right]^{2}eV(t)\cos\phi_{s}(t)}$$
(3-6b)

The form of Eq.(3-5) reminds us of the betatron oscillations in a

transport line with a nonlinear component. In analogy to perturbed betatron oscillations, we separate the right-hand side of Eq.(3-5) into unperturbed and perturbing terms

$$K^{(0)}(x, p; Y) = \frac{1}{2}p^2 + \frac{1}{2}\lambda_{\bullet}(Y)x^2,$$
 (3-7a)

$$K^{(1)}(x, p; \gamma) = -\frac{1}{3} \lambda_1(\gamma) x^3.$$
 (3-7b)

In the next section, we shall discuss the phase dynamics in the vicinity of transition by studying the above linear Hamiltonian  $K^{(o)}$ .

# § 4. Linear Motion

We consider the linear system described by

$$K^{(p)}(x, p; \gamma) = \frac{1}{2} [p^2 + \lambda_0(\gamma) x^2].$$
 (4-1)

We make the assumption that the peak RF voltage and the synchronous phase angle are constant near transition. At the transition energy, the quantity

$$\alpha^{(0)} - 1/T_s^2(t),$$

vanishes. Then, in the vicinity of this energy this quantity can be approximated by the first term in a Tayler series expansion of Eq.(3-6a) for deviations of  $\Upsilon$  from  $\Upsilon_{\tau}$ . We can therefore write

$$\lambda_o(\tau) \simeq - \frac{4\pi h \left[ \Upsilon_s(\tau) - \mathcal{E}_{\tau} \right]}{\beta_s^2(0) F_s(0) eV(0) (os\phi_s(0) \mathcal{E}_{\tau}^3)}. \tag{4-2}$$

Further, the quantity  $\gamma_{\bf S}(t) - \gamma_{\bf T}$  can be written in terms of

$$Y_{s}(t) - Y_{T} = Y_{s} t$$

$$= \frac{eVSi'n\Phi_{s}}{m_{0}c^{2}T_{s}} t$$

$$= -\frac{Si'n\Phi_{s}(0)}{m_{0}c^{2}(os\Phi_{s}(0))} \gamma. \qquad (4-3)$$

Substitution of Eq.(4-3) into Eq.(4-2) yields

$$\lambda_0(\tau) \simeq \frac{4\pi h \sin \phi_s}{\beta_s^2 E_s^2 \delta_1^2 eV \cos^2 \phi_s} \gamma$$
 (4-4)

Here all quantities are evaluated at transition. For the sake of later simplicity we set

$$R^2 = \frac{4\pi h \sin \phi_3}{\beta_s^2 E_s^2 \Gamma_T^2 e V \cos^2 \phi_s}.$$

We point out that the system represented by Eq.(4-1) has an exact dynamical invariant, which is designated as the "Courant-Snyder invariant" /8,9/

$$T(x, p; \tau) = \frac{1}{2S(\tau)} \left\{ x^2 + \left[ \frac{1}{2} \dot{s}(\tau) x - S(\tau) p \right]^2 \right\}, (4-5)$$

where  $S(\Upsilon)$  satisfies the auxiliary differential equation

$$\frac{1}{2}S\dot{S} - \frac{1}{4}\dot{S}^2 + \lambda_0(\gamma)S^2 = 1. \tag{4-6}$$

When  $\lambda_{ullet}(\gamma)$  is constant, the invariant I is identical to the action variable of the system if we take the initial condition

$$S(+\infty) = 1/\sqrt{\lambda_0(\infty)} , S(+\infty) = 0.$$
 (4-7)

In the following,  $S(\gamma)$  will be called a synchrotron amplitude function. For a time-varying function  $\lambda_{\bullet}(\gamma)$ , from Eq.(4-5) we

know that an infinite sequence of phase points that have a certain constant value of I at an arbitrary time behaves as a deformable moving ellipse in the phase space  $(x,p;\gamma)$  after that time. The form of such a ellipse, called an "invariant curve" in the following, is uniquely determined by the auxiliary differential equation (4-6) alone. We consider the invariant curve described in terms of

$$I(\alpha, \beta; \alpha) = I_o$$
 (4-8)

with constant  $I_0$ . The quantity I is equal to the value of the action variable of the infinite set of phase points that comprise the invariant curve, as mentioned above.

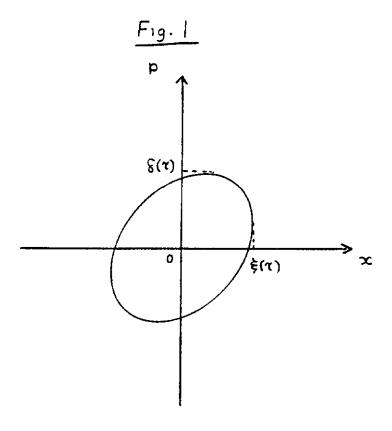
We may characterize the ellipse by two parameters  $\xi$  ,  $\delta$  which are functions of  $S(\Upsilon)$  and  $\dot{S}(\Upsilon)$ 

$$\xi(\tau) = \sqrt{2I_{o}S(\tau)}, \qquad (4-9a)$$

$$\xi(\tau) = \sqrt{2I_o \frac{\left[1 + \dot{s}(\tau)/4\right]}{S(\tau)}}.$$
 (4-9b)

They are the maximum extent of the ellipse in x and  $\rho$ , respectively.

If we assume that all parameters change adiabatically after  $\gamma = \gamma_1$ , we can choose the approximated initial conditions



$$S(\tau_i) = 1/\sqrt{\lambda_o(\tau_i)} , S(\tau_i) = 0. \qquad (4-10)$$

At  $\gamma = \gamma_1$ , the ellipse begins to move, following the time evolution of  $S(\gamma)$  which is determined by Eq.(4-6). If we know the values of  $S(\gamma)$ ,  $\dot{S}(\gamma)$  at  $\gamma = 0$ , which is the transition time, we can evaluate exactly the maximum upper or lower height from the synchronous energy  $\xi(0)$  and the half phase spread S(0).

For the present  $\lambda_{\rm e}(\Upsilon)$ , we know the general solution of Eq.(4-6) can be written in the terms of Bessel functions (see Appendix A),

$$S(\tau) = \left(\frac{\pi}{3}\right)^{2} \gamma \left[a N_{y_{3}}^{2}(z) + b J_{y_{3}}^{2}(z) - 2 \left(ab - \frac{q}{\pi^{2}}\right)^{2} J_{y_{3}}(z) N_{y_{3}}(z)\right], \quad (4-11)$$

where a and b are arbitrary coefficients that must be determined by the initial conditions (4-10), and

$$Z = \frac{2}{3} \notin \gamma^{3/2}.$$

After mathematical manipulation (see Appendix B), we have the coefficients

$$a = \frac{3}{2} z_1 J_{y_3}^{2}(z_1) + \frac{3}{2 z_1 N_{y_3}^{2}(z_1)} \left[ z_1 J_{z_3}(z_1) N_{y_3}(z_1) + \frac{z}{\pi} \right]^{2} - \frac{6}{\pi z_1 N_{y_3}(z_1)} \left[ z_1 J_{z_3}(z_1) N_{y_3}(z_1) + \frac{1}{\pi} \right],$$
 (4-12a)

$$b = \frac{3}{2} z_1 N_{y_3}^2(z_1) + \frac{3}{2z_1 J_{y_1}^2(z_1)} \left[ z_1 J_{y_3}^2(z_1) N_{y_3}(z_1) + \frac{2}{\pi} \right]^2 \quad (4-12b)$$

where

$$Z_i = \frac{2}{3} \pm \gamma_i^{3/2}$$
 (4-13)

The coefficients a and b have been uniquely determined by the initial conditions and we now know the exact time evolution of the invariant curve. In particular, we are interested in the invariant curve just at transition; it represents a bunch envelope. From Eq.(4-11) we obtain the values of  $S(\Upsilon)$  and  $S(\Upsilon)$  at  $\Upsilon=0$  (see Appendix C)

$$S(0) = \frac{4}{3} \left(\frac{\pi}{3}\right)^2 \left(\frac{\pi}{3}\right)^{-2/3} \frac{a}{\int_{-2}^{2} (2/3)}$$
 (4-14a)

$$\dot{S}(0) = \frac{2\pi}{3} \left[ -\frac{a}{\sqrt{3}} + (ab - \frac{q}{\pi^2})^{\frac{1}{2}} \right]. \tag{4-14b}$$

Introduction of Eqs.(4-14a),(4-14b) into Eqs.(4-4a),(4-4b) leads to analytical expressions for the maximum upper or lower height from the synchronous energy and the half spread around the synchronous phase at transition, that is,

$$\xi(o) = \xi(x_i) \sqrt{s(o)/s(x_i)}, \qquad (4-15a)$$

$$S(o) = S(\gamma_1) \sqrt{\left[1 + \frac{s(o)}{4}\right] s(\gamma_1) / S(o)}. \tag{4-15b}$$

We consider the case with the initial conditions at  $\Upsilon_i = +\infty$  where the linear  $\Upsilon$ -dependence of  $\lambda_0(\Upsilon)$  still hold to good approximation. At such a region, the Bessel and Neumann functions become sinusoidal with equal amplitude and quadrature phase relationship, namely,

$$J_{y_3}(z_1) = \sqrt{2/\pi z_1} \cos(z_1 - 5\pi/12),$$

$$N_{y_3}(z_1) = \sqrt{2/\pi z_1} \sin(z_1 - 5\pi/12),$$
also
$$J_{-2/3}(z_1) = \sqrt{2/\pi z_1} \cos(z_1 + \pi/12)$$

$$= -\sqrt{2/\pi z_1} \sin(z_1 - 5\pi/12).$$
(4-16)

Substitution of Eq.(4-16) into Eqs.(4-12a) and (4-12b) yields

$$a = b = 3/\pi$$
. (4-17)

From Eq.(4-17), we can obtain a universal relationship between  $\xi$  and  $\delta$ 

$$\xi(0) \delta(0) = \xi(\tau_1) \delta(\tau_1) \left[ 1 + \frac{\dot{s}(0)^2}{4} \right]^{1/2}$$

$$= \frac{2}{\sqrt{3}} \xi(\tau_1) \delta(\tau_1). \tag{4-18}$$

Equation (4-18) is equivalent to the result obtained by Hereward /11/.

Predictions of the linear theory are not discussed in detail here. Nevertheless they are in very good agreement with numerical simulations. These simulations have been performed following the exact mapping equations (2-1a) and (2-1b) where nonlinear kinematic terms are not included in order to verify the validity of the linear theory. In addition, the linear theory discussed here which provides exact time evolution of bunch shapes can re-establish the well-known story /8/ associated with transition crossing.

# § 5. Nonlinear Motion

The nonlinear kinematic term has been distinguished as a perturbing term in Eq.(3-7b). It gives unignorable effects to the liear oscillation only during a very short period when a particle crosses the transition energy. In order to assess its quantitative effects, it is convenient to use the action-angle formalism.

Under the linear canonical transformation

$$Q = g^{-1}x,$$
  
 $P = -\dot{g}x + gp,$  (5-1)

where  $\Im(\tau) = \sqrt{\Im(\tau)}$  satisfies the auxiliary equation

$$9 + \lambda_0(\tau) 9 = 9^{-3}$$
, (5-2)

the Hamiltonian (3-5) reduces to

$$K(Q,P;\tau) = g^{-2} \left[ \frac{1}{2} P^2 + \frac{1}{2} Q^2 - \frac{g^5}{3} \lambda_1(\tau) Q^3 \right].$$
 (5-3)

If a change of independent variable

$$\theta(\tau) = \int_{\gamma_2}^{\gamma} s^{-2}(\gamma) d\gamma + \Theta_0, \qquad (5-4)$$

is made, the Hamiltonian (5-3) becomes

$$K'(Q,P;\theta) = \frac{1}{2}(P^2 + Q^2) - \frac{\rho^5}{3}\lambda_1(r)Q^3$$
 (5-5)

Furthermore, introduction of the action-angle variables  $(\psi, \mathbf{J})$ 

$$Q = \sqrt{2J} \sin \psi,$$

$$P = \sqrt{2J} \cos \psi,$$
(5-6)

yields the Hamiltonian

$$G(\Psi, \mathcal{I}; \theta) = \mathcal{I} - \frac{p^5}{3} \lambda_i(\tau) (2\mathcal{I})^3 \sin^3 \Psi.$$
 (5-7)

From Eq.(5-7), we derive the canonical equations

$$\Psi' = \frac{\partial G}{\partial J} = 1 - 9^5 \lambda_1(\tau) (zJ)^{1/2} \sin^3 \Psi,$$
 (5-8a)

$$J' = -\frac{\partial G}{\partial \Psi} = 9^5 \lambda_1(x) (2)^{3/2} \sin^2 \Psi \cos \Psi.$$
 (5-8b)

If the perturbing term in Eq.(5-8a) is much less than the unperturbed term, that is, small compared with unit, we obtain to first order

$$\Im(\theta(r)) \simeq \Im(\theta(\omega)) + \left[2\Im(\theta(\omega))\right]^{3/2} \int_{\theta(\omega)}^{5/2} \int_{\theta(\omega)}^{5/2} \gamma^{5}(\tau) \lambda_{1}(\tau) \sin^{2}\theta(\tau) \cos^{2}\theta(\tau) \frac{d\theta_{2}}{(5-9)}$$

Using the relation (5-4), we have a more convenient integral expression

$$J(r) \simeq J(r_2) + \left[2J(r_2)\right]^{3/2} \int_{r_2}^{r} e^3(r) \lambda_1(r) \sin^2\theta(r) \cos\theta(r) dr \qquad (5-10)$$

where the lower boundary of the integration will be determined in the following considerations.

We note that the period of phase oscillation has the same notation as the synchrotron amplitude function  $S(\Upsilon)$  because the oscillation frequency is described by

$$v = \psi(s) = \frac{1}{s(r)}. \tag{5-11}$$

From Eq.(5-8b), the typical modulation period of the perturbing term is  $S(\Upsilon)/2$ . Generally, the value of  $S(\Upsilon_i)/2$  is much smaller than  $\Upsilon_i$ . This fact means that the effects of the perturbing term are averaged out, at least, at the early stage of non-adiabatic motion ( $0 \ll \Upsilon \leq \Upsilon_i$ ). So it is reasonable for us to take S(0)/2 as the typical lower boundary  $\Upsilon_2$  when the perturbation begins to retain net effects. Fortunately, in many real situations,  $\Upsilon_2$  is sufficiently small so as to satisfy the condition

$$\vec{x}(\tau_2) = \frac{2}{3} R \tau_2^{3/2} \ll 1.$$
 (5-12)

The relation (5-12) enables us to describe  $P(\Upsilon)$  or  $S(\Upsilon)$  by

elementary functions. In order to do this, it is necessary to know the Bessel function for small value of z. In Ref.13 we observe that

$$J_{1/3}(z) = \frac{1}{\Gamma(4/3)} \left(\frac{z}{2}\right)^{1/3},$$
 (5-13a)

$$N_{1/3}(z) = \frac{1}{\sin \pi/3} \left[ \frac{\cos \pi/3}{\Gamma(4/3)} \left( \frac{z}{2} \right)^{1/3} - \frac{1}{\Gamma(2/3)} \left( \frac{z}{2} \right)^{1/3} \right]. \quad (5-13b)$$

Considering Eqs.(5-13a) and (5-13b), we obtain to first order with respect to  $\boldsymbol{\gamma}$ 

$$S(\gamma) = \varphi^{2}(\gamma) = \left(\frac{\pi}{3}\right)^{2} \left(-\gamma \gamma + r\right)$$
 (5-14)

where

$$\mathcal{F} = \frac{49}{3\Gamma(4/3)\Gamma(2/3)},$$
 (5-15a)

$$r = \frac{4a}{3\Gamma^{2}(2/3)} \left(\frac{R}{3}\right)^{2/3} + \frac{4}{\sqrt{3}} \left(ab - \frac{9}{\pi^{2}}\right)^{1/2} \frac{1}{\Gamma(2/3)\Gamma(4/3)}. \quad (5-15b)$$

Further substitution of the above result (5-14) into (5-4) yields

$$\theta(\tau) = \int_{\gamma_2}^{\gamma_2} 1/s(\gamma') d\gamma' + \theta_0$$

$$= \left(\frac{3}{\pi}\right)^2 \left(-\frac{1}{4}\right) \left[\log\left(\frac{\pi}{3}\right)^2 \left(-\eta\gamma + r\right) - \log\left(\frac{\pi}{3}\right)^2 \left(-\eta\gamma + r\right)\right] + \theta_0.$$
(5-16)

On the other hand,  $\lambda_i(\tau)$  can be expanded with respect to  $\tau_s - \tau_{\tau}$  and to first order

$$\lambda_{1}(\tau) = -\frac{2\pi h \left[d_{n} + 3\beta_{s}^{2}/2\right]}{\beta_{s}^{4} \left(m_{o}c^{2}\right)^{2} eV \cos\phi_{s} \delta_{T}^{4}} \left[1 + \frac{4\tan\phi_{s}}{m_{o}c^{2} \Upsilon_{T}} \tau\right], \quad (5-17)$$

where  $\alpha_n(\epsilon_0^{(i)}/\alpha^{(i)})$  is the nonlinear lattice parameter. Using Eqs.(5-14) and (5-17), we write the perturbing integration in Eq.(5-10) in terms of

$$\Delta J = \left[ 2J(r_{1}) \right]^{\frac{3}{2}} \lambda_{1}(0) \int_{r_{2}}^{\sqrt{2}} \left( 1 - \frac{r_{1}}{r_{1}} \tau^{\frac{3}{2}} \right) \left( 1 + \frac{4 \tan \phi_{3}}{moc^{2} \delta_{1}} \tau \right) \sin^{2}\theta h_{1}(0) \theta h_{2}(0) d \tau$$

$$= \left[ 2J(r_{1}) \right]^{\frac{3}{2}} \lambda_{1}(0) \int_{r_{1}}^{\sqrt{2}} \left[ 1 + \left( \frac{4 \tan \phi_{3}}{moc^{2} \kappa_{1}} - \frac{3 r_{1}}{2 r_{1}} \right) \tau \right] \sin^{2}\theta h_{2}(0) \theta h_{3}(0) d \tau.$$

Here we set

$$g_1 = \frac{4 \tan \Phi_s}{m_0 c^2 V_T} - \frac{3 R}{2 r}$$
, (5-19)

$$\theta(x) = \frac{9}{3} \log \left( \frac{\gamma}{g_2} + g_3 \right) + \overline{\theta_0}, \qquad (5-20)$$

$$g_2 = \left(\frac{3}{\pi}\right)^2 \left(-\frac{1}{4}\right), \quad g_3 = S(0) = \left(\frac{\pi}{3}\right)^2 r, \quad \overline{\theta}_0 = -g_1 \log \left(\frac{\gamma_2}{g_2} + g_3\right) + \theta_0.$$

Using the identity

$$\sin^2\theta(\tau)\cos\theta(\tau) = \frac{1}{4}\left[\cos\theta(\tau) - \cos3\theta(\tau)\right],$$

with together (5-19), (5-20), we have

$$\Delta J \simeq \frac{1}{4} \left[ 2J(\tau_1) \int_{0}^{3/2} \lambda_{1}(0) \int_{0}^{3/2} (0) \int_{0}^{\infty} (1+g_{1}\tau) \times \left[ \cos \left( \frac{\gamma}{g_{1}} + g_{3} \right) + \overline{\theta_{0}} \right] - \cos 3 \left[ g_{1} \log \left( \frac{\gamma}{g_{1}} + g_{3} \right) + \overline{\theta_{0}} \right] d\tau_{(5-21)}$$

We are interested in the value of  $\Delta J$  at  $\Upsilon$  =0. This value is evaluated in the following way.

The integration to be performed is

$$I = \int_{\gamma_{1}}^{0} d\gamma (1+g_{1}\gamma) \left\{ \cos \left( g_{1} \log \left( \frac{\gamma}{g_{1}} + g_{3} \right) + \overline{\theta}_{0} \right\} - (05) \left[ g_{1} \log \left( \frac{\gamma}{g_{2}} + g_{3} \right) + \overline{\theta}_{0} \right] \right\} = 0.53 \left[ g_{1} \log \left( \frac{\gamma}{g_{2}} + g_{3} \right) + \overline{\theta}_{0} \right]$$

Change of the integration variable to

$$w = g_2 \log \left( \frac{\tau}{g_2} + g_3 \right) + \overline{\theta}_0,$$

leads to

$$I = e^{\frac{\overline{\theta}_{0}}{g_{2}}} \int_{w_{1}}^{w_{2}} dw e^{\frac{w}{g_{1}}} \left(1 - g_{1}g_{2}g_{3} + g_{1}g_{2}e^{-\frac{\overline{\theta}_{0}}{g_{1}}}e^{w/g_{1}}\right) \times (5-23)$$

$$w_1 = g_2 \log \left( \frac{\gamma_2}{g_2} + g_3 \right) + \overline{g_0} = \theta_0,$$
 (5-24a)

$$w^2 = g_2 \log g_3 + \overline{\theta}_0 = g_2 \log \left[g_3 / (\frac{\gamma_2}{2} + g_3)\right] + \theta_0$$
 (5-24b)

The integration of Eq.(5-23) is trivial. We obtain

$$I = e^{-\frac{\pi}{2}\sqrt{3}2} \left( \left( 1 - g_1 g_2 g_3 \right) \right) \frac{1}{1/g_2^2 + 1} \left| e^{w/g_2} \left( \frac{1}{g_2} (osw + sinw) \right|_{w_1}^{w_2} \right) - \frac{1}{1/g_2^2 + 1} \left| e^{w/g_2} \left( \frac{1}{g_2} (osw + 3sinw) \right|_{w_1}^{w_2} \right)$$

$$g_1 g_2 e^{-\frac{\pi}{2}\sqrt{3}2} \left\{ \frac{1}{4/g_2^2 + 1} \left| e^{2w/g_3} \left( \frac{2}{g_2} (osw + sinw) \right) \right|_{w_1}^{w_2} \right\} - \frac{1}{4/g_3^2 + 9} \left[ e^{2w/g_3} \left( \frac{2}{g_2} (osw + 3sinw) \right) \right]_{w_1}^{w_2} \right\} = - \frac{1}{4/g_3^2 + 9} \left[ e^{2w/g_3} \left( \frac{2}{g_2} (osw + 3sinw) \right) \right]_{w_1}^{w_2} \right\} = - \frac{1}{4/g_3^2 + 9} \left[ e^{2w/g_3} \left( \frac{2}{g_2} (osw + 3sinw) \right) \right]_{w_1}^{w_2} \right] = - \frac{1}{4/g_3^2 + 9} \left[ e^{2w/g_3} \left( \frac{2}{g_2} (osw + 3sinw) \right) \right]_{w_3}^{w_2} \right] = - \frac{1}{4/g_3^2 + 9} \left[ e^{2w/g_3} \left( \frac{2}{g_2} (osw + 3sinw) \right) \right]_{w_3}^{w_3} \right] = - \frac{1}{4/g_3^2 + 9} \left[ e^{2w/g_3} \left( \frac{2}{g_2} (osw + 3sinw) \right) \right]_{w_3}^{w_3} \right] = - \frac{1}{4/g_3^2 + 9} \left[ e^{2w/g_3} \left( \frac{2}{g_2} (osw + 3sinw) \right) \right]_{w_3}^{w_3} \right] = - \frac{1}{4/g_3^2 + 9} \left[ e^{2w/g_3} \left( \frac{2}{g_2} (osw + 3sinw) \right) \right]_{w_3}^{w_3} \right] = - \frac{1}{4/g_3^2 + 9} \left[ e^{2w/g_3} \left( \frac{2}{g_2} (osw + 3sinw) \right) \right]_{w_3}^{w_3} \right] = - \frac{1}{4/g_3^2 + 9} \left[ e^{2w/g_3} \left( \frac{2}{g_2} (osw + 3sinw) \right) \right]_{w_3}^{w_3} \right] = - \frac{1}{4/g_3^2 + 9} \left[ e^{2w/g_3} \left( \frac{2}{g_2} (osw + 3sinw) \right) \right]_{w_3}^{w_3} \right] = - \frac{1}{4/g_3^2 + 9} \left[ e^{2w/g_3} \left( \frac{2}{g_2} (osw + 3sinw) \right) \right]_{w_3}^{w_3} \right] = - \frac{1}{4/g_3^2 + 9} \left[ e^{2w/g_3} \left( \frac{2}{g_3} (osw + 3sinw) \right) \right]_{w_3}^{w_3} = - \frac{1}{4/g_3^2 + 9} \left[ e^{2w/g_3} \left( \frac{2}{g_3} (osw + 3sinw) \right) \right]_{w_3}^{w_3} \right]$$

Some of the intermediate steps in the calculation are explained in Appendix D. The final results are

$$I = \frac{9.93(1-9.3.3)}{1+92^2} \left\{ \left[ -\sin\theta + 92\cos\theta - 92(1/292+1) \right] \sin\theta + \left[ \cos\theta + 92\sin\theta - (1/292+1) \right] \cos\theta \right\}$$

$$- \frac{9.93(1-9.3.3)}{1+932^2} \left\{ \left[ -\sin3\theta + 392\cos3\theta - 392(1/292+1) \right] \sin3\theta + \left[ \cos3\theta + 392\sin3\theta - (1/292+1) \right] \cos3\theta \right\}$$

$$+ \left[ \cos3\theta + 392\sin3\theta - (1/292+1) \right] \cos3\theta \right\}$$

$$+ \frac{9.9293^2}{4+92^2} \left\{ \left[ -2\sin\theta + 92\cos\theta - 92(1/292+1) \right] \sin\theta + \left[ 2\cos\theta + 92\sin\theta - 2(1/292+1)^2 \right] \cos\theta \right\}$$

$$- \frac{9.9293^2}{4+932^2} \left\{ \left[ -2\sin\theta + 392\cos3\theta - 392(1/292+1)^2 \right] \sin\theta \right\}$$

$$+ \left[ 2\cos\theta + 392\cos3\theta - 392(1/292+1)^2 \right] \sin3\theta$$

$$+ \left[ 2\cos\theta + 392\cos3\theta - 2(1/292+1)^2 \right] \cos3\theta$$

$$+ \left[ 2\cos\theta + 392\cos3\theta - 2(1/292+1)^2 \right] \cos3\theta$$

$$\Theta = g_2 \log \left[ 1 / \left( \frac{1}{2g_2} + 1 \right) \right].$$
(5-27)

The value of I is dependent of the initial phase  $\theta_{\bullet}$ . If the maximum and minimum values of I are denoted by Imax(>0) and Imin(<0), we can write the positive and negative changes of the action variable in terms of

$$\Delta J(0)_{\pm} = \frac{\left[2J(7_2)\right]^{3/2}}{4} \lambda_1(0) S^{3/2}(0) I_{\text{Min}}^{\text{max}}, \qquad (5-28)$$

where the suffix symbols of the left-hand side do not always correspond to those of the right-hand side because of the negative sign of  $\lambda_1(0)$ . We define the emittance increase parameter, which is independent of the initial emittance (the initial value of the action variable) of a particle, by

$$\mathcal{H}_{\pm} = \frac{1}{2} \lambda_1(0) S^{2}(0) I_{\text{Min}}$$
(5-29)

It is assumed that there are no net effects of the perturbing term up to  $\Upsilon = \Upsilon_2$  and we can take the value of  $J(\Upsilon_1)$  as  $J(\Upsilon_2)$ . Then

$$J(x_i) \left[ 1 + K_{-} \left[ 2J(x_i) \right] \le J(0) \le J(x_i) \left[ 1 + K_{+} \left[ 2J(x_i) \right] \right]$$
(5-30)

Thus incoherent changes of the action variable give an unsymmetric

ellipse in the phase space. In particular, we can obtain expressions for the maximum upper and lower height from the synchronous energy and the maximum and minimum excursion from the synchronous phase as

$$\xi(0)_{\text{min}}^{\text{max}} = \left\{2J(z_1)S(0)[1+k(\pm\sqrt{2J(z_1)})] = \xi(z_1)\frac{S(0)}{S(z_1)}[1+k(\pm\sqrt{2J(z_1)})]^{\frac{1}{2}}\right\}$$
(5-31a)

$$\delta(0) \max_{\min} = \delta(\tau_1) \sqrt{\frac{s(\tau_1)}{s(0)}} \left[ 1 + k_{\pm} \sqrt{2J(\tau_1)} \right]^{\frac{1}{2}}$$
 (5-31b)

where the value of  $S(0)/S(Y_i)$  has already been derived in the previous section.

# § 6. Applications

The theoretical considerations are applied to an example that corresponds to the nominal acceleration mode in the FNAL Main Ring /12/. For this example, several parameters of acceleration are listed in Table 1.

Table 1 Acceleration Parameters of the FNAL Main Ring

Harmonics h=1113RF Voltage eV=-2.(MV)Synchronous phase  $f_s=235.87^{\circ}$ Transition gamma  $f_s=18.8$ 

Transition energy E =17.639(GeV)

Nonlinear lattice parameter d/d =0.14 (Ref.11)

We choose Y, as

$$\gamma_1 = (os \phi_s(o) (E_s(o) - E_s(\gamma_1))$$
  
= 559.9428 MeV, (6-1)

where  $E_{S}(\Upsilon_{i})$  is enough far from transition. Using the parameters in the table and the value for  $\Upsilon_{i}$  , we get

$$R = \left[ \frac{4\pi h \left| \sin \phi_s \right|}{E_s^2(0) \delta_T^2 \left( os^2 \phi_s |eV| \right)} \right]^{\frac{1}{2}} = 4.089 |x| 0^{-4} \text{ MeV}, \quad (6-2)$$

then

$$z_1 = \frac{2}{3} \pi \gamma_1^{3/2} = 3.6120$$
  
 $S(\gamma_1) = 1/\pi \gamma_1^{1/2} = 103.3476 \text{ MeV}$  (6-3)

From a table of Bessel functions /14/, we read

$$J_{3}(z_1) = -0.2736$$
,  $N_{3}(z_1) = 0.3172$ ,  $J_{23}(z_1) = -0.309$ . (6-4)

Substitution of these values for  $z_1$ ,  $J_{ij}(z_1)$ ,  $N_{ij}(z_1)$ , and  $J_{ij}(z_1)$  into Eqs.(4-12a) and (4-12b) gives

$$a = 0.4056 + 0.3296 + 0.1877 = 0.9229,$$

$$b = 0.5451 + 0.443 = 0.9881.$$
(6-5)

From Eqs.(6-2) and (6-5) we get the value of the synchrotron amplitude function at transition

$$S(0) = \frac{4}{3} \left(\frac{\pi}{3}\right)^2 \left(\frac{\pi}{3}\right)^{-2/3} \frac{q}{\Gamma^2(2/3)} = 277.503 \text{ (MeV)}. (5-6)$$

Here, we use

$$\Gamma(1/3) = 2.6801$$
,  $\Gamma(2/3) = 1.3550$ ,  $\Gamma(4/3) = 0.89338 /15/$ 

The value of the coefficient  $\lambda_i(0)$  is

$$|\lambda_{1}(0)| = \frac{2\pi h \left[\alpha^{(1)}/\alpha^{(0)} + 3\beta_{s}^{2}(0)/2\right]}{\delta_{1}^{4}\beta_{s}^{4}(0) \left(\omega_{0}c^{2}\right)^{2} eV(0) \left(\cos\phi_{s}(0)\right)} = 1.2343 \times 10^{-7} \text{ MeV}_{(6-7)}^{-3}$$

The S(0) above leads to the value of the typical boundary

$$\gamma_2 = S(0)/2 = 138.7515 \text{ (MeV)},$$
 (6-8)

The parameters  $g_i$ ,  $g_2$  in the integration of Eq.(5-18) are

$$g_2 = -0.8279/a = -0.8971,$$

$$g_1 = 3.6118 * 10^{-4} - \frac{3}{2} \frac{\Gamma(2/3)}{\Gamma(4/3)} \left(\frac{1}{3}\right)^{2/3} = -5.66 * 10^{-3}$$

then

$$\Theta = g_2 \log \left[ 1 / \left( \frac{1}{2g_2} + 1 \right) \right] = -0.7312 (6-10)$$

Substituting these values for  $g_1$ ,  $g_2$ ,  $g_3$ ,  $\Theta$  into (5-26), we obtain the integral I in the form

$$I = (-39.1614) * \sin(\Theta_0) + (-72.9502) * \cos(\Theta_0) + (71.9840) * \sin(3\Theta_0) + (-5.2259) * \cos(3\Theta_0) + (6-11)$$

Thus 
$$Imax = 96$$
,  $Imin = -174$ . (6-12)

From Eq.(6-12), we have

$$\aleph_{+} = 1/2*0.9577*10^{-7}*4622.7619*179 = 3.85*10^{-2}$$
(6-13)

Finally, substituting these values for S(0), S(7) into Eq.(5-31a), we get

$$\xi(0) = \xi(x_i) *1.6386 * [1. + 3.7871*10^{-3} \xi(x_i)]_{i}^{i}$$

$$(6-14)$$

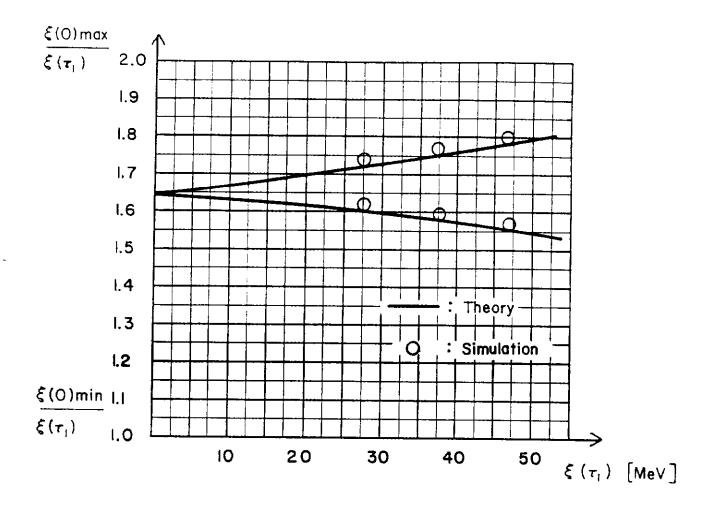
$$\xi(0) = \xi(x_i) *1.6386 * [1. - 2.0894*10^{-3} \xi(x_i)]_{i}^{i/2}$$

For several initial emittances which still allow linear approximations, results obtained from Eq.(6-14) are plotted in Fig.2. In the same figure, numerical simulation results are also given. We see quite good agreement.

# § 7. Emittance Blowup at Transition

When a particle crosses the transition energy, the electric phase of RF is abruptly changed externally to  $\pi$  -  $\phi_s$ . Such a manipulation yields time-reversal of the phase motion for the dynamical system described by

$$H(x, p; \gamma) = \frac{1}{2} \left[ p^2 + \lambda_0(\gamma) x^2 \right], \qquad (7-1)$$



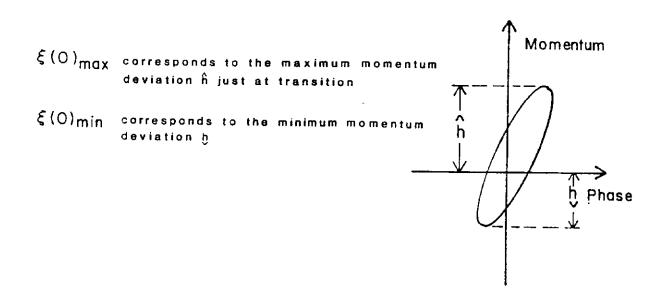
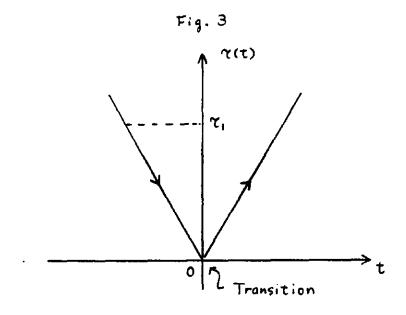


FIGURE 2.

because of the change of sign of the cosine function in

$$\gamma(t) = \int_{0}^{t} -\left[\frac{eV(t)(os\phi_{s}|t)}{T_{s}(t)}\right]dt, \qquad (7-2)$$

will lead to time-reversing as seen in Fig. 3



while the sign of  $\lambda_o(\Upsilon)$ ,

$$\lambda_o(\tau) = -\frac{4\pi h \left[ r_s(t) - r_T \right]}{\beta_s^2 \text{ (7-3)}}$$

still remains unchanged due to sign changes of the denominator and numerator. This holds even if all higher order terms with respect to phase are included. Consequently, emittance blowup during transition crossing can't in principle be explained by synchrotron oscillation theory, which restricts itself to ordinary pendulum oscillations with adiabatically changing coefficients.

On the other hand, the coefficient of the nonlinear kinematic term.

$$\lambda_{i}(\tau) = -\frac{2\pi h \left[\alpha^{(i)} + 3\beta_{s}^{2}(t)/2\delta_{s}^{3}(t)\right]}{\left[L\beta_{s}(t)E_{s}(t)\right]^{2}eV(t)\left(\cos\varphi_{s}(t)\right)}, \quad (7-4)$$

changes its sign after the phase-jump. The dynamical system including such a term is therefore no longer timereversible. In other words, synchrotron oscillations accompanied with kinematic nonlinearity are singular at transition. Just after passing transition, a bunch suddenly meets an unmatched bucket. This leads to actual emittance blowup. The magnitude of the blowup is proportional to the amount of emittance distortion due to the nonlinear kinematic term. Thus we may write the final emittance blowup ratio during transition crossing as

$$R = 1 + 2 \cdot K_{+} \cdot \sqrt{2J(r_{1})},$$
 (7-5)

where  $J(\Upsilon_i)$  is the emittance or the value of the action variable far below transition. The final blowup ratio R is plotted as a function of the initial emittance for the normal acceleration mode (h=1113) in Fig.4. Results of measurements are also given in the figure. The small overestimate seen may imply that the exact nonlinear lattice parameter is somewhat smaller than the value used in the present calculations.

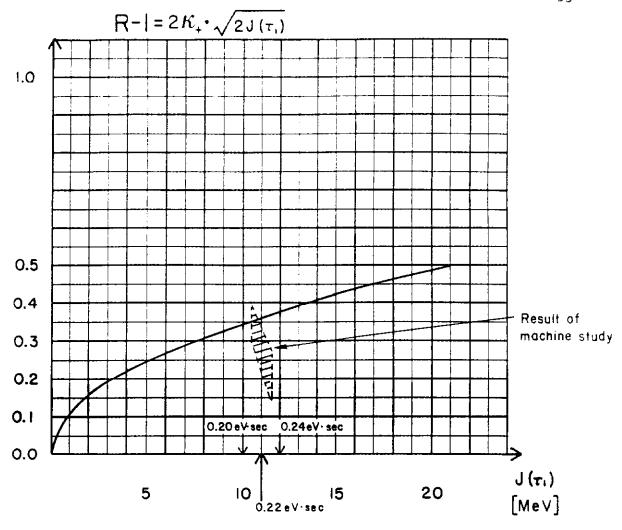


FIGURE 4. EMITTANCE BLOW-UP RATIO

## § 8. Conclusion

A linear classical theory of transition, which is equivalent to the usual one of matrix form,/11,16/ has been developed by introducing the synchrotron amplitude function. As a natural extension of this linear theory, a general method which relies on perturbation techniqes to assess effects of the lowest order nonlinear kinematic term is presented. When these theories are applied to the case of the FNAL Main Ring, they agree very well with results of computer simulations and real measurements. This emphasizes the importance of the higher-order chromaticity control, which can be done by adjusting the n-th and 2n-th Fourier sextupole magnetic fields (n ≥ horizontal components of the betatron tune) /4/. If $\lambda_{i}(\Upsilon)$  is reduced by such higher order chromaticity control, the emittance blowup discussed here will be improved.

The present analytical approach can be also used to derive explicit expressions for emittance increments resulting from other nonlinear forces which become significant, in particular, in the vicinity of transition.

## Acknowledgments

The author wishes to thank Dr. J.E.Griffin for stimulation to his studying the present topic and for his offer of valuable data of machine studies. He wishes to thank Dr. S.Ohnuma for useful discussions and Dr. F.T.Cole for careful reading of the manuscript. He also would like to thank the head of the Tevatron I Section, Dr. J.Peoples, for his interest in this work and for his continual encouragement.

Appendix A General Solution of Auxiliary Equation

When  $x_1(\Upsilon)$  and  $x_2(\Upsilon)$  are linearly independent solutions of the time dependent linear equation

$$\frac{\partial}{\partial x} + \lambda_0(x) = 0, \qquad (A-1)$$

we can write a general solution of its modified nonlinear auxiliary equation (or envelope equation)

$$g + \lambda_0(x) g = \frac{1}{g_3}, \qquad (A-2)$$

where  $\Upsilon(\Upsilon)$  is the square root function of  $S(\Upsilon)$ , in terms of  $x_1(\Upsilon)$  and  $x_2(\Upsilon)$  as

$$P(x) = \left( C_1 x_1^2 + C_2 x_2^2 + C_3 x_1 x_2 \right)^{1/2}. \tag{A-3}$$

Squaring both sides of (A-3) and Differentiating with respect to the independent variable  $\gamma$  , we have

$$2 \dot{\beta}^{2} + 2 \dot{\beta} \dot{\beta} = 2 \dot{C}_{1} \dot{x}_{1}^{2} + 2 \dot{C}_{1} \dot{x}_{1} \dot{x}_{1} + 2 \dot{C}_{2} \dot{x}_{2}^{2} + 2 \dot{C}_{2} \dot{x}_{2} \dot{x}_{2}$$

$$+ \dot{C}_{3} \dot{x}_{1} \dot{x}_{2} + \dot{C}_{3} \dot{x}_{1} \dot{x}_{2} + 2 \dot{C}_{3} \dot{x}_{1} \dot{x}_{2}, \qquad (A-4)$$

From (A-1) and (A-2), Eq.(A-4) reduces to an equation including first time derivatives alone. Further, using (A-3), we obtain

$$LHS = \frac{(2C_1X_1\dot{\chi}_1 + 2C_2X_2\dot{\chi}_2 + C_3X_1\chi_2 + C_3X_1\dot{\chi}_2)^2}{2(C_1X_1^2 + C_2X_2^2 + C_3X_1\chi_2)} + \frac{2}{C_1X_1^2 + C_2X_2^2 + C_3X_1\chi_2}$$

$$- 2\lambda(\Upsilon)(C_1X_1^2 + C_2X_2^2 + C_3X_1\chi_2),$$

RH\$ =  $2C_1\dot{x_1}^2 + 2C_2\dot{x_2}^2 + 2C_3\dot{x_1}\dot{x_2} - 2\lambda(\chi)(C_1\chi_1^2 + C_2\chi_2^2 + C_3\chi_1\chi_2)$ Equating both sides and eliminating terms, we find

$$(C_3^2 - 4C_1C_2)(\dot{x_1}x_2 - x_1\dot{x_2})^2 + 4 = 0.$$
 (A-5)

From (A-5), the arbitrary constant  $C_1$ ,  $C_2$ , and  $C_3$  are not independent. Namely,  $C_3$  is determined from  $C_1$  and  $C_2$  as

$$C_3 = \pm \sqrt{4C_1C_2 - \frac{4}{w^2}}$$
, (A-6)

where W is the Wronskian  $x_1 \dot{x_2} - \dot{x_1} x_2$ , which is a constant. Consequently, we can write the general solution of (4-6) as

$$S(\tau) = C_1 \chi_1^2(\tau) + C_2 \chi_2^2(\tau) - 2 \sqrt{C_1 C_2 - \frac{1}{W^2}} \chi_1(\tau) \chi_2(\tau). \quad (A-7)$$

For the present case

$$\lambda(\tau) = \pi^2 \gamma , \qquad (A-8)$$

the independent solutions of the linear equation can be written in terms of Bessel and Neumann functions of order 1/3,

$$x_1(r) = \gamma^{1/2} N_{1/3} \left( \frac{2}{3} R \gamma^{3/2} \right),$$
 (A-9a)

$$\mathcal{X}_{\bullet}(\Upsilon) = \Upsilon^{1/2} \left( \frac{2}{3} R \Upsilon^{3/4} \right). \tag{A-9b}$$

Thus the general solution becomes

$$S(\tau) = \gamma \left[ a N_{y_3}^2 \left( \frac{2}{3} R \gamma^{3/2} \right) + b J_{y_3}^2 \left( \frac{2}{3} R \gamma^{3/2} \right) - 2 \sqrt{ab - \frac{1}{w^2}} x \right]$$

$$N_{y_3} \left( \frac{2}{3} R \gamma^{3/2} \right) J_{y_3} \left$$

with

$$W = x_1(0) x_2(0) - x_1(0) x_2(0) = 3/\pi c.$$

This is in agreement with the result obtained by Lewis /10/.

Appendix B Particular Solution of the Auxiliary Equation (4-6)

Since the initial conditions are

$$S(r_1) = 1/\frac{r_1^{1/2}}{r_1}$$
 (B-1a)

$$\dot{S}(r_i) = 0, \tag{B-1b}$$

we have immediately two algebraic equations for a and b

$$4N_{\nu}^{2} + bJ_{\nu}^{2} - 2(ab - \frac{9}{\pi^{2}})^{1/2}J_{\nu}N_{\nu} = \frac{6}{\pi^{2}E_{1}},$$
 (B-2a)

$$aN_{\nu}N_{\nu}' + bJ_{\nu}J_{\nu}' - (ab - \frac{q}{\pi^{2}})^{1/2}(J_{\nu}'N_{\nu} + J_{\nu}N_{\nu}') = -\frac{2}{\pi^{2}E_{i}^{2}},$$
 (B-2b)

where  $\nu$  is 1/3, the prime denotes derivatives with respect to z, and all Bessel functions are to be given their values at z1=z( $\chi_1$ ). From (3-2a), we have

$$(ab - \frac{q}{\pi^2})^{1/2} = \frac{aNv^2 + bJ_v^2 - 6/\pi^2z_1}{2J_vN_v}.$$
 (B-3)

Substitution of (B-3) into (B-2b) yields

$$\frac{N_{\nu}(N_{\nu}J_{\nu} - N_{\nu}J_{\nu}')}{2J_{\nu}} a + \frac{J_{\nu}(J_{\nu}'N_{\nu} - J_{\nu}N_{\nu}')}{2N_{\nu}} b = \frac{3(J_{\nu}'N_{\nu} + J_{\nu}N_{\nu}')}{\pi^{2}Z_{\nu}J_{\nu}N_{\nu}} \frac{2}{\pi^{2}Z_{\nu}J_{\nu}N_{\nu}} \frac{2}{\pi^{2}Z_{\nu}J_{\nu}N_{\nu}}$$

Using the formula

$$J_{\nu}N_{\nu}' - J_{\nu}'N_{\nu} = 2/\pi \mathcal{B}, \qquad (B-5)$$

we have

$$Q = \frac{J_0^2}{N_0^2} b - \frac{3}{\pi N_0^2} \left( J_0' N_0 + J_0 N_0' + \frac{2 J_0 N_0}{3 Z_1} \right). \tag{B-6}$$

If we substitute (B+6) into (B+3), square both sides, and compare corresponding terms, we obtain

$$b = \frac{3}{2} z_1 N_{\nu}^2 + \frac{1}{6 z_1 J_{\nu}^2} \left[ \frac{3}{2} (J_{\nu}' N_{\nu} + J_{\nu} N_{\nu}') + J_{\nu} N_{\nu} + \frac{3}{\pi c} \right]^2.$$
 (B-7)

Further, from a recursion equation for Bessel functions

$$J_{\nu}' = J_{\nu-1} - \nu z^{-1} J_{\nu}, \qquad (B-8)$$

and the relation (B-5), we find

$$J_{\nu}'N_{\nu} + J_{\nu}N_{\nu}' = J_{\nu}'N_{\nu} + J_{\nu}'N_{\nu} + \frac{2}{\pi z_{1}}$$

$$= 2(J_{\nu} - \frac{\nu}{z_{1}}J_{\nu})N_{\nu} + \frac{2}{\pi z_{1}}.$$
(B-9)

Introduction of (B-9) into (B-7) leads to

$$b = \frac{3}{2} z_1 N_0^2 + \frac{3}{2 z_1 J_0^2} \left( z_1 J_0 N_0 + \frac{2}{\pi} \right)^2.$$
 (B-10)

Thus we also find

$$\alpha = \frac{3}{2}z_{1}J_{0}^{2} + \frac{3}{2z_{1}N_{0}^{2}}(z_{1}J_{0}N_{0} + \frac{2}{\pi})^{2} - \frac{6(z_{1}J_{0}N_{0} + 1/\pi)}{\pi z_{1}N_{0}^{2}}.$$
 (B-11)

Appendix C Evaluation of S(0) and  $\dot{S}(0)$ 

We set

$$S(0) = \lim_{\tau \to 0} \left(\frac{\pi}{3}\right)^{2} \left[a N_{y_{3}}^{2}(z(\tau)) + b J_{y_{3}}^{2}(z(\tau)) - 2(ab - \frac{9}{\pi^{2}})^{2} J_{y_{3}}(z(\tau)) N_{y_{3}}(z(\tau))\right]^{(C-1)}$$

where  $z(\gamma)$  is  $\frac{2}{3} \, \xi \, \gamma^{3/2}$  . Using the formula

$$Ny_3(z) = \frac{1}{\sin \pi/3} \left[ \cos \frac{\pi}{3} \cdot J_{y_3}(z) - J_{-y_3}(z) \right],$$
 (C-2)

S(0) becomes

$$\left(\frac{\pi}{3}\right)^{2} \lim_{n \to \infty} \gamma \left\{ \frac{q}{3} \left[ J_{N_{3}}^{2} - 4J_{N_{3}} J_{-N_{3}} + 4J_{-N_{3}}^{2} \right] + bJ_{N_{3}}^{2} - \frac{2}{\sqrt{3}} \left(ab - \frac{q}{\pi^{2}}\right)^{1/2} \left(J_{N_{3}}^{2} - 2J_{N_{3}} J_{N_{3}}\right)^{2} \left(C - 3\right) \right\}$$

If we retain only the lowest order term of the series expansion of the Bessel functions, we have

$$J_{V_3}(\bar{z}(\gamma)) = \left(\frac{1}{3}\right)^3 \frac{\gamma^{1/2}}{\Gamma(4/3)}, \qquad (C-4a)$$

$$J_{-1/3}(3(x)) = \left(\frac{\pi}{3}\right)^{-1/3} \frac{\chi^{-1/2}}{\Gamma(2/3)}.$$
 (C-4b)

Substituting (C-4a) and (C-4b) into (C-3) and taking its limit at  $\gamma$  =0, we obtain

$$S(6) = \frac{4}{3} \left(\frac{\pi}{3}\right)^{2} \left(\frac{\pi}{3}\right)^{-2/3} \frac{1}{\int_{-1}^{2} \frac{1}{(2/3)}}.$$
 (C-5)

S (0) can also be evaluated in a similar way. We set

$$\dot{S}(0) = \lim_{\chi \to 0} \left( \frac{\pi}{3} \right)^{2} \left[ a N y_{3}^{2} + b J y_{3}^{2} - 2 \left( ab - \frac{q}{\pi^{2}} \right)^{2} J y_{3} N y_{3} \right]$$

$$+ 2 \chi \dot{z} \left[ a N y_{3} N y_{3} + b J y_{3} J y_{3}^{2} - \left( ab - \frac{q}{\pi^{2}} \right)^{2} \left( T y_{1}^{2} N y_{2} + J y_{3} N y_{3}^{2} \right) \right] , \qquad (C-6)$$

where dots denote derivatives with respect to  $\boldsymbol{\gamma}$ , primes denote derivatives with respect to z, and all Bessel functions are functions of  $z(\boldsymbol{\gamma})$ . Using formulas (B-10), (C-2), and the recursion relation (B-9), we have

$$N_{13}N_{13}^{1} = \frac{1}{3} \left( J_{13} - 2J_{-1/3} \right) \left( J_{-2/3} + 2J_{-1/3} - \frac{J_{1/3}}{32} + \frac{2}{32} J_{-1/3} \right), \qquad (C-7a)$$

$$J_{13}J_{13}' = J_{13}' \left(J_{-2/3} - \frac{1}{32}J_{13}\right),$$
(C-7b)

$$J_{3}'N_{3} + J_{3}N_{3}' = \frac{2}{\sqrt{3}}(J_{43} - \frac{1}{32}J_{y_{3}})(J_{y_{3}} - 2J_{43}) + \frac{2}{\pi 2}.$$
(C-7c)

Retaining only the lowest-order term of the series expansion for  $J_{\frac{2}{3}}(z(\Upsilon))$  and  $J_{\frac{2}{3}}(z(\Upsilon))$ , we have

$$J_{\mathcal{H}}\left(z\left(z\right)\right) = \left(\frac{z}{3}\right)^{\frac{1}{3}} \frac{\gamma}{\Gamma(5/3)}, \qquad (C-8a)$$

$$J_{2/3}(2(\gamma)) = \left(\frac{\pi}{3}\right)^{2/3} \frac{\gamma^{-1}}{\Gamma(1/3)}.$$
 (2-8b)

From  $\gamma z = k \gamma^{3/2}$ , the limiting values of the component terms at

$$\lim_{\gamma \to 0} \gamma \stackrel{?}{=} J_{1/3} J_{2/3} = \lim_{\gamma \to 0} \frac{1}{R} \left(\frac{\frac{R}{3}}{3}\right)^{\frac{1}{3} - \frac{2}{3}} \frac{\gamma h_{2} + h_{2} - 1}{P(4/3)P(1/3)} = 0,$$

$$\lim_{\gamma \to 0} \gamma \stackrel{?}{=} J_{1/3} J_{2/3} = \lim_{\gamma \to 0} \frac{1}{R} \left(\frac{\frac{R}{3}}{3}\right)^{\frac{1}{3} + \frac{2}{3}} \frac{\gamma h_{2} + h_{2} + 1}{P(4/3)P(1/3)} = 0,$$

$$\lim_{\gamma \to 0} \gamma \stackrel{?}{=} J_{1/3} J_{2/3} = \lim_{\gamma \to 0} \frac{1}{R} \left(\frac{\frac{R}{3}}{3}\right)^{\frac{1}{3} + \frac{2}{3}} \frac{\gamma h_{2} + h_{2} + 1}{P(4/3)P(1/3)} = \frac{3}{P(2/3)P(1/3)},$$

$$\lim_{\gamma \to 0} \gamma \stackrel{?}{=} J_{1/3} J_{2/3} = \lim_{\gamma \to 0} \frac{1}{R} \left(\frac{\frac{R}{3}}{3}\right)^{\frac{1}{3} + \frac{2}{3}} \frac{2^{\frac{1}{3} + \frac{1}{3}} + \frac{2}{3}}{P(2/3)P(1/3)} = 0,$$

$$\lim_{\gamma \to 0} \gamma \stackrel{?}{=} J_{1/3} J_{2/3} = \lim_{\gamma \to 0} \frac{1}{R} \left(\frac{\frac{R}{3}}{3}\right)^{\frac{1}{3} + \frac{1}{3}} \frac{2^{\frac{1}{3} + \frac{1}{3}} + \frac{2}{3}}{P(2/3)P(1/3)} = 0,$$

$$\lim_{\gamma \to 0} \gamma \stackrel{?}{=} J_{1/3} J_{2/3} = \lim_{\gamma \to 0} \frac{1}{R} \left(\frac{\frac{R}{3}}{3}\right)^{\frac{1}{3} + \frac{1}{3}} \frac{2^{\frac{1}{3} + \frac{1}{3}} + \frac{2}{3}}{P(2/3)P(1/3)} = 0,$$

$$\lim_{\gamma \to 0} \gamma \stackrel{?}{=} J_{1/3} J_{2/3} = \lim_{\gamma \to 0} \frac{1}{R} \left(\frac{\frac{R}{3}}{3}\right)^{\frac{1}{3} + \frac{1}{3}} \frac{2^{\frac{1}{3} + \frac{1}{3}} + \frac{1}{2}}{P(2/3)P(1/3)} = 0,$$

$$\lim_{\gamma \to 0} \gamma \stackrel{?}{=} J_{1/3} J_{2/3} = \lim_{\gamma \to 0} \frac{1}{R} \left(\frac{\frac{R}{3}}{3}\right)^{\frac{1}{3} + \frac{1}{3}} \frac{2^{\frac{1}{3} + \frac{1}{3}} + \frac{1}{2}}{P(2/3)P(1/3)} = 0.$$

$$\lim_{\gamma \to 0} \gamma \stackrel{?}{=} J_{1/3} J_{2/3} = \lim_{\gamma \to 0} \frac{1}{R} \left(\frac{\frac{R}{3}}{3}\right)^{\frac{1}{3} + \frac{1}{3}} \frac{2^{\frac{1}{3} + \frac{1}{3}} + \frac{1}{2}}{P(2/3)P(1/3)} = 0.$$

$$\lim_{\gamma \to 0} \gamma \stackrel{?}{=} J_{1/3} J_{2/3} = \lim_{\gamma \to 0} \frac{1}{R} \left(\frac{\frac{R}{3}}{3}\right)^{\frac{1}{3} + \frac{1}{3}} \frac{2^{\frac{1}{3} + \frac{1}{3}} + \frac{1}{2}}{P(2/3)P(1/3)} = 0.$$

Substitution of (C-9) into (C-6) yields

$$\dot{S}(0) = \left(\frac{\pi}{3}\right)^{2} \left\{ \lim_{\chi \to 0} \left\{ \left[ q \left( \frac{4}{3} \right] \right]_{1/3}^{2} - \frac{8\tau \dot{z}}{qz} \right]_{-1/3}^{2} \right\} + a \left( -\frac{4}{3} \int_{1/3}^{1/3} + \frac{8\tau \dot{z}}{qz} \int_{1/3}^{1/3} \right) \\
+ \left( ab - \frac{q}{\pi^{2}} \right)^{1/2} \left( \frac{4}{\sqrt{3}} \int_{1/3}^{1/3} - \frac{8\tau \dot{z}}{3\sqrt{3}z} \int_{1/3}^{1/3} \int_{-1/3}^{1/3} \right) \right] \\
+ \left[ -\frac{4a}{\Gamma(2/3)\Gamma(1/3)} - 2\left( ab - \frac{q}{\pi^{2}} \right)^{1/2} \left( -\frac{12}{\sqrt{3}\Gamma(2/3)\Gamma(1/3)} + \frac{2\tau \dot{z}}{\pi z} \right) \right] \right\}.$$
(C-10)

Thus, we have

$$\dot{S}(0) = \frac{2\pi}{3} \left[ -\frac{q}{\sqrt{3}} + \left(ab - \frac{q}{\pi^2}\right)^{1/2} \right].$$
 (C-11)

Here we use the relation

$$\Gamma(2/3) \Gamma(1/3) = \frac{2\pi}{\sqrt{3}}$$
.

Appendix D Calculation of Perturbing Integration

One of the four parts in the perturbation integration,

$$I_{A} = e^{-\overline{\theta}d\theta_{2}} \left\langle \frac{1 - 9.9293}{1/92 + 1} \middle| e^{w/92} \left( \frac{1}{92} (osw + sinw) \middle|_{w_{1}}^{w_{2}} \right), \quad (D-1)$$

is calculated as follows:

$$I_{A} = \frac{g_{2}(1-g_{1}g_{2}g_{3})}{1+g_{2}^{2}} e^{-\theta d_{3}} \left[ e^{w_{2}/g_{2}} (\cos w_{1} + g_{2}\sin w_{2}) - e^{\frac{\pi}{3}} (\cos w_{1} + g_{2}\sin w_{2}) \right]$$

$$= \frac{g_{2}(1-g_{1}g_{2}g_{3})}{1+g_{2}^{2}} \left[ e^{\log g_{3}} (\cos w_{1} + g_{2}\sin w_{2}) - e^{\log g_{3}(\frac{1}{2g_{2}}+1)} (\cos w_{1} + g_{2}\sin w_{1}) \right]$$

$$= \frac{g_{2}(1-g_{1}g_{2}g_{3})}{1+g_{2}^{2}} \left\{ g_{3}(\cos \left[g_{2}\log\left(\frac{1}{|y_{2}|+1} + \theta_{0}\right)\right] + g_{2}g_{3}\sin\left[g_{2}\log\left(\frac{1}{|y_{2}|+1} + \theta_{0}\right)\right] - g_{3}\left(\frac{1}{|y_{2}|+1} + \theta_{0}\right) \right\},$$

$$T_{A} = \frac{g_{2}g_{3}(1-g_{1}g_{2}g_{3})}{1+g_{2}^{2}} \left\{ \left[ \sin \Theta + g_{1}\cos \Theta - g_{2}(\frac{1}{2g_{2}}+1) \right] \sin \Theta \right\},$$

$$(D-2)$$
where

The other three parts can be calculated in a similar way.

## References

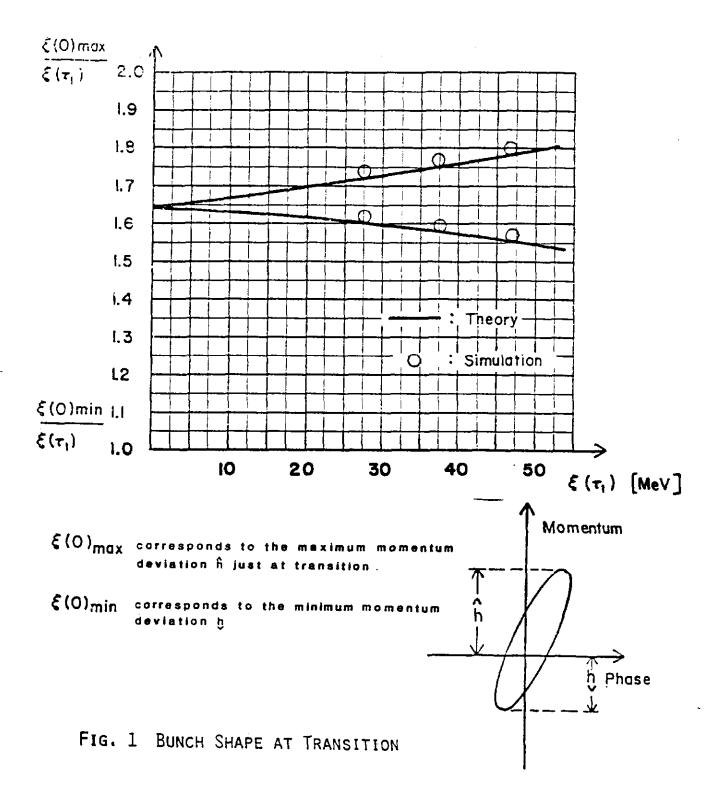
- /1/ K.R.Symon and A.M.Sessler, Proc. of the CERN Symp. on the High Energy Accelerators, Vol.1, (1956) p44.
- /2/ K.Johnsen, Proc. of the CERN Symp. on the High-Energy Accelerators, Vol.1, (1956) p106.
- /3/ H.Bruck and et al., IEEE Trans. Nucl. Sci. NS-20, (1973) p822.
- /4/ A.Ando and K.Takayama, to be published in Proc. of 1983 Particle Accelerators Conference, (1983), or FNAL Report TM-1073, (1981).
- /5/ J.E.Griffin and et al., to be published in Proc. of 1983 Particle

  Accelerators Conferences, (1983).
- /6/ A.Sorenssen, Particle Accelerators, Vol. 6, (1975) p141.
- /7/ A.Sorenssen and H.G.Hereward, CERN Report MPS/Int. MU/EP66-1 (1966).
- /8/ E.D.Courant and H.S.Snyder, Ann. Phys. 3 (1958) pl.
- /9/ K.Takayama, to be published in Proc. of 1983 Particle Accelerators

  Conference, (1983), or FNAL Report FN-349 (1981).
- /10/H.R.Lewis, J. of Math. Phys., Vol.9, No.11, (1968) p1983.
- /11/H.G.Hereward, CERN Report MPS/DL Int.66-3, (1966).
- /12/For example, J.Griffin, AIP Conf. Proc. No.87 (1982) p564.
- /13/S.Ohnuma, private communication, (1982).
- /14/G.N.Watson, in: Theory of Bessel Functions (Cambridge at the University Press, 1966) p714.
- /15/E. Jahnke & et al, in: Tables of Higher Functions (McGraw-Hill Book Company, New York, 1960), p176.
- /16/J.C.Herrera, Particle Accelerators, Vol.3, (1972) p49.

## Abstract

The phase dynamics of small amplitude synchrotron oscillations in the vicinity of the transition energy is discussed with kinematic nonlinearities included. We introduce a synchrotron amplitude function analogous to the betatron amplitude function and solve analytically the time evolution of bunch shapes, where the kinematic nonlinearities result in unsymmetric bunch shapes. In addition, the above synchrotron oscillation is singular at transition crossing because of the kinematic nonlinearity. From this simple fact, we identify an inherent source of bunch diffusion. A method for estimating its size is presented. When this theory is applied to the case of the FNAL Main Ring, the predictions are in good agreement with numerical simulations and are not inconsistent with experimental results.



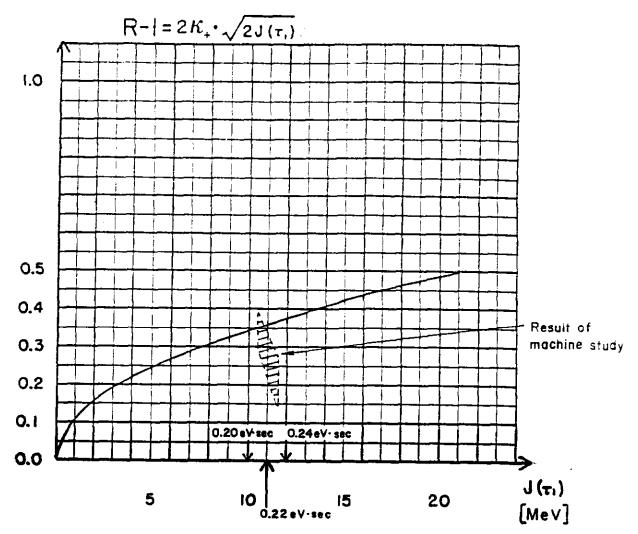


FIG. 2 EMITTANCE BLOWUP RATIO